

The new Hamiltonian has now been found (to the order ε) and so has the determining function, where

$$S_1 = \int F_{1p} dl = (v - l), \quad (\text{J.7.15})$$

the equation of the center. With

$$F^*(L^*, G^*) = \frac{\mu^2}{2L^{*2}} + \varepsilon \frac{\mu^2}{L^{*3}G^*}, \quad (\text{J.7.16})$$

Hamilton's equations have the solution $L^* = \text{const.}$, $G^* = \text{const.}$,

$$l^* = l_0 + \left(\frac{\mu^2}{L^{*3}} + \varepsilon \frac{3\mu^2}{L^{*4}G^*} \right) t, \quad g^* = g_0 + \varepsilon \frac{\mu^2}{L^{*3}G^{*2}} t, \quad (\text{J.7.17})$$

giving the secular perturbations. Then, from

$$S = L^*l + G^*g + \varepsilon(v - l) \quad (\text{J.7.18})$$

we have

$$L = L^* + \varepsilon \left(\frac{\partial v}{\partial l} - 1 \right), \quad G = G^*, \quad l^* = l + \varepsilon \frac{\partial v}{\partial L}, \quad g^* = g + \frac{\partial v}{\partial G}, \quad (\text{J.7.19})$$

to the first order in ε . $\frac{\partial v}{\partial L}$ can be found at once from the integral expression, (J.7.16), for S_1 . The remaining derivatives can be found by applying the chain rule using formulas for elliptic motion. This is done in Ref. 19, where the approach just used, is applied to the problem of the motion of an artificial satellite. This method was proposed by Poincaré, and applied to orbits of minor planets by von Zeipel; it is often referred to as the "von Zeipel method."

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