

Optical Depth & Albedo of
Ejecta from
DEEP IMPACT

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The goal

- To extract some physical properties of the cometary nuclei from DEEP IMPACT images of 9P/Tempel 1.

Optical Depth & Albedo

$$I/I_0 = e^{-\tau} \quad \text{where:}$$

τ = optical depth

I = intensity

For simple scattering and absorption:

$$\tau = N \sigma \quad \text{where:}$$

σ = extinction cross section

N = column density

Optical Depth & Albedo

- A single particle (grain of dust or whatever) scatters light anisotropically.
- For a single particle: $I = F_{\text{sun}} \sigma p f(\theta)$, where:
 - p = geometric albedo
 - $f(\theta)$ = scattering function (θ = phase angle)
 - F_{sun} = solar flux received by the comet

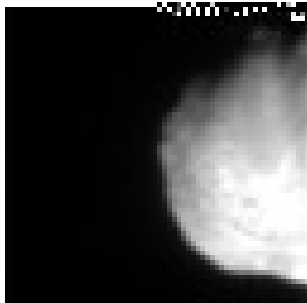
Method

- We use the approximation that the optical depth is not large so that $I(\text{column}) = N * I(\text{particle})$. Then we have:

$$I = F_{\text{sun}} \tau p f(\theta)$$

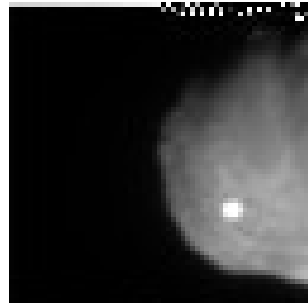
The Data

- Deep Impact – MRI VIS Calibrated Images
9P/Tempel 1 Encounter Phase
DOY 2005-185 (2005-07-04)



$t = -1.288 \text{ s}$

(mv0173727701_9000910_043_rr.fit)



$t = +0.165 \text{ s}$

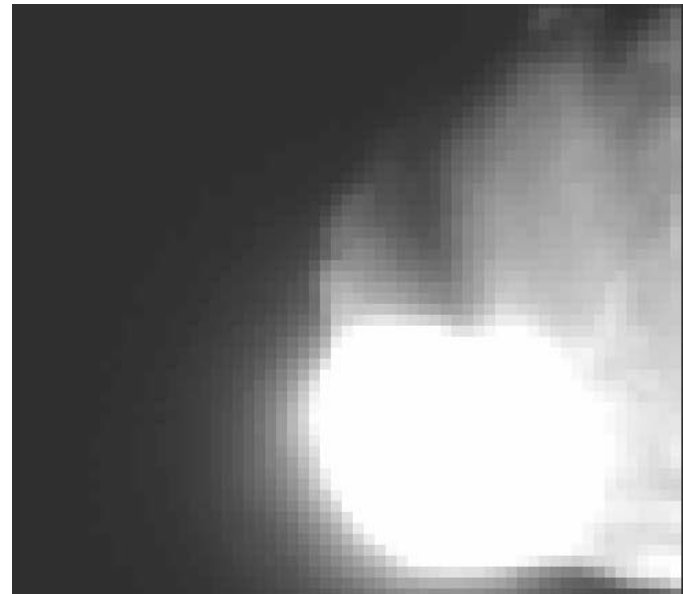
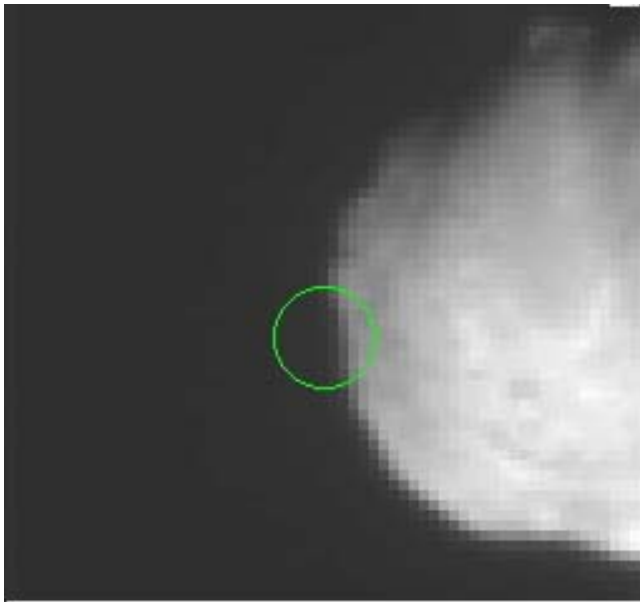
(mv0173727703_9000910_066_rr.fit)



$t = +1.454 \text{ s}$

(mv017372704_9000910_043_rr.fit)

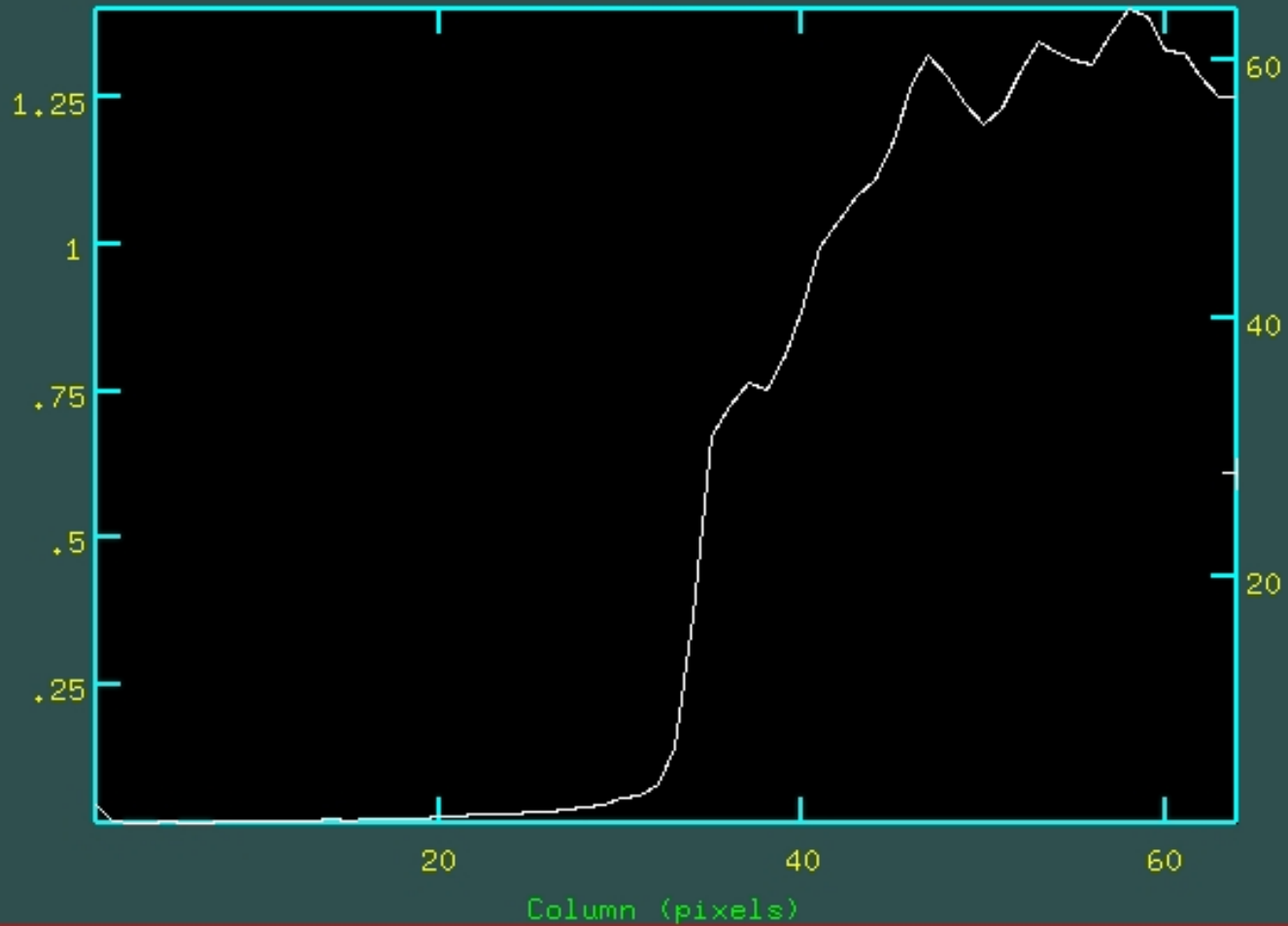
Making the measurements



We selected a feature on the sunward side of the impact site, close to the limb (better contrast).

irafterm

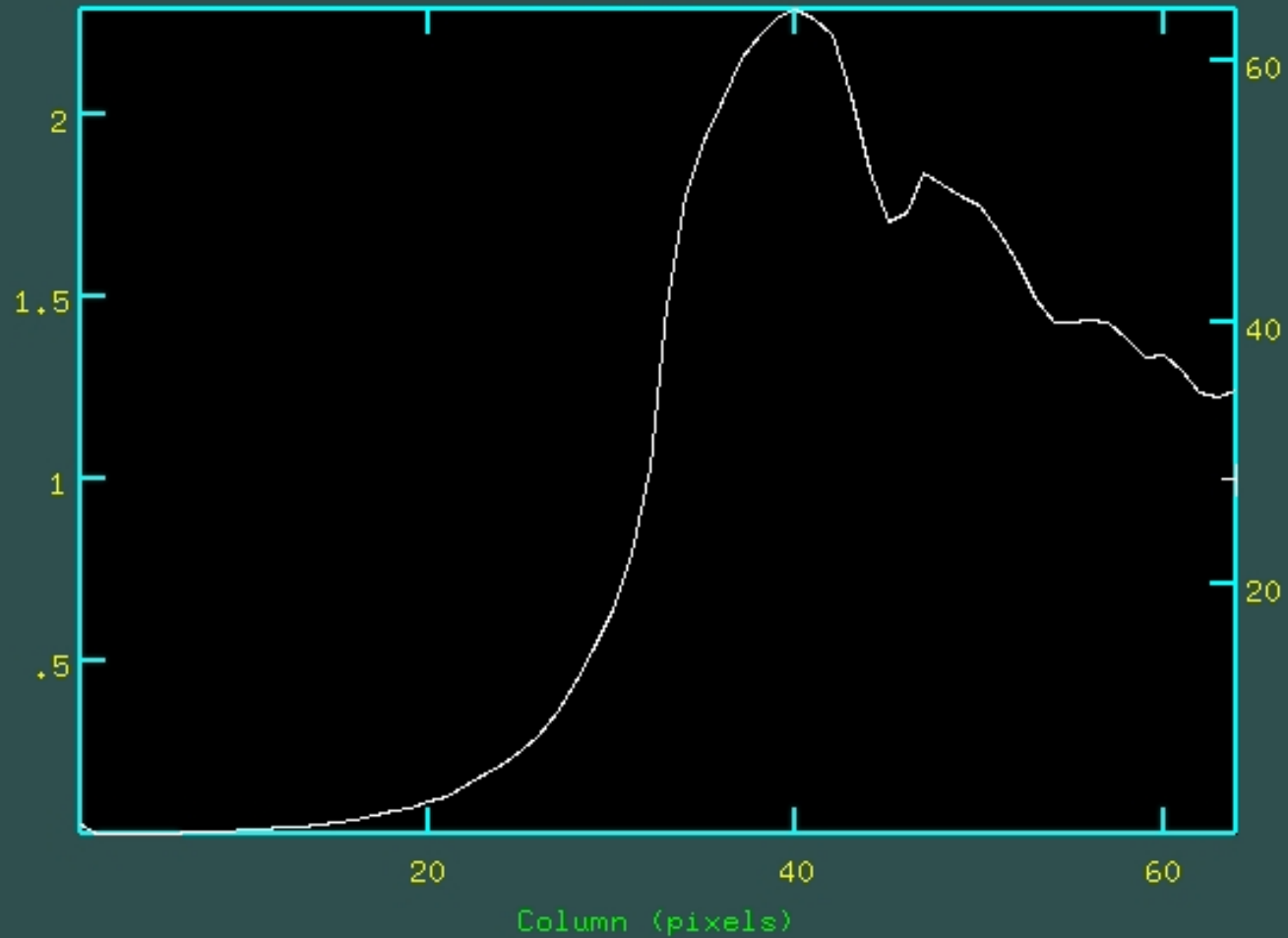
NOAO/IRAF V2.12.2a-EXPORT estudiantes@auladoc1 Thu 07:37:46 02-Aug-2007
Line 28 of img1[0].fit
9P/TEMPEL 1



pixel=[32,28] value=0.07940182

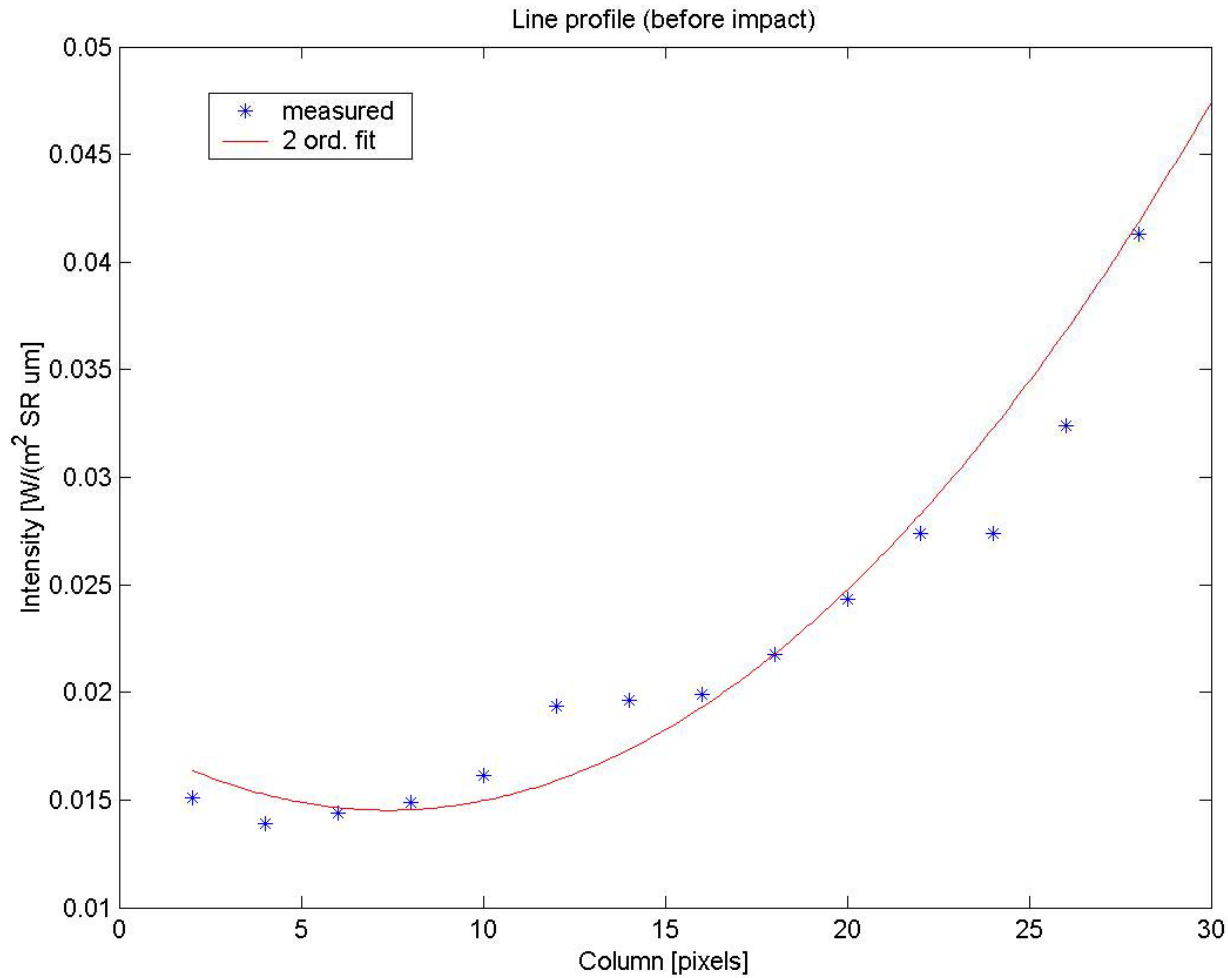
irafterm

NOAO/IRAF V2.12.2a-EXPORT estudiantes@auladoc1 Thu 06:57:16 02-Aug-2007
Line 28 of img2[0].fit
9P/TEMPEL 1



pixel=[30,28] value=0.642195

Extrapolating the coma brightness (I_0)



Very Preliminary Results

- We made a rough estimate of the optical depth and then of the albedo of the ejecta.
- Values used:
 - $\theta = 62^\circ$, from the label file
 - $f(\theta) = 0.0595$, from Lumme-Bowell law
 - $F_{\text{sun}} = \frac{L_{\text{sun}}}{4\pi r^2}$

(heliocentric distance $r = 1.5$ UA, from the label file)

Very Preliminary Results

$$\tau = \log(I/I_0) = 2.45 \quad \longrightarrow \quad P = \frac{I}{F_{\text{sun}} \tau f(\theta)} = 0.01$$

The value obtained is lower than the standard value for cometary nuclei ($P = 0.04$).

To be continued...

- Improve the technique by adjusting the scale and offset before register the images (to make sure that we make all measurements at the same feature on all images).
- Extending the project by analyzing the variation of albedo with position around the limb and with time.

To be continued...

- This can show variations in the type of particles (ice *vs.* dust and organics) both with direction of ejection and with depth of excavation.